Hybrid Approach to Broadband Ground Motion Simulations

Graves and Pitarka (2010, 2015)

- Semi-deterministic approach at low frequencies
- Semi-stochastic approach at high frequencies
- Kinematic Rupture Generator
 - Unified scaling rules for rise time, rupture speed and corner frequency
 - Depth scaling of rise time (increase) and rupture speed (decrease)
 required to model shallow (< 5 km) moment release
- Validation studies using recorded ground motions from California and Eastern US earthquakes



Hybrid Simulation Method

Low Frequency (f < 1 Hz)

- Complete kinematic finite-fault rupture description including spatial and temporal heterogeneity in the slip function.
- Full theoretical Green's functions computed for specified plane-layered or 3D velocity structure including anealstic attenuation.

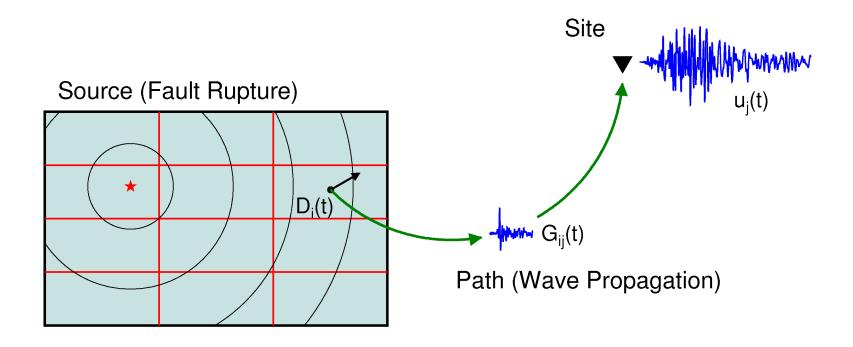
High Frequency (f > 1 Hz)

- Limited kinematic finite-fault rupture description including spatial heterogeneity in slip and rupture time.
- Each subfault radiates an ω^{-2} spectrum with **stochastic** phase and conically averaged radiation pattern.
- Simplified ray path Green's functions include travel time and impedance effects.
- Frequency dependent attenuation represents both anelastic and scattering effects.



Representation Theorem

$$u_i(t) = \sum_i G_{ii}(t) * D_i(t)$$





Low Frequency Representation Theorem

The low frequency simulation utilizes the basic representation theorem presented earlier

$$u_{j}(t) = \sum_{i=1,N} G_{ij}(t) * D_{i}(t)$$

where the $D_i(t)$ are given by the rupture characterization described in the previous section and the $G_{ii}(t)$ are Green's functions (GFs) describing wave propagation from the i^{th} subfault to the j^{th} receiver.

The current implementation of the Broadband Platform is restricted to 1D velocity structures. The use of 1D media allows for very efficient use and storage of the GFs.

In our implementation, the GFs are computed using the frequency-wavenumber (FK) technique of Zhu and Rivera (2002). The FK GFs contain the full theoretical waveform response from zero frequency to the upper limit specified in the calculation (typically several Hz). The Zhu and Rivera method computes GFs for 3 fundamental fault orientations, from which any arbitrary faulting mechanism can be computed using a linear combination of the fundamental fault responses.

The FK computation itself is not currently installed on the BBP. For the GP method, we pre-compute a database of GFs for a specified 1D velocity model and then this GF database is installed on the BBP.

For efficiency, the GFs are computed for a matrix of depths and distances covering the anticipated range of these parameters that might be encountered in the simulations. For WUS GFs, the depth range is 0-30 km and the distance range is 0-500 km. For ENA, the depth range is 0-35 km and the distance range is 0-1100 km.



High Frequency Representation Theorem

The high frequency simulation approach is based on the "stochastic" method first introduced for point sources by Boore (1993). The extension to finite-faults is described by Frankel (1995), Beresnev and Atkinson (1997), and Hartzell et al. (1999), among many others. The representation is constructed in the frequency domain and aims to match the ω^2 amplitude spectrum as described by Brune (1970).

Denoting A(f) as the Fourier transform of the ground acceleration waveform a(t) observed at a particular site for a specified fault rupture, this can then be represented as the summation of the individual responses A_i(f) from each subfault, where N is the total number of subfaults.

$$\mathcal{F}$$
 {a(t)} = A(f) = $\sum_{i=1,N} A_i(f)$

The acceleration spectrum for the ith subfault is given by

$$A_{i}(f) = \sum_{j=1,M} C_{ij} S_{i}(f) G_{ij}(f) P(f) W_{i}^{*}(f)$$

where the summation j=1,M accounts for different possible ray paths (e.g., direct, Moho-reflected), and

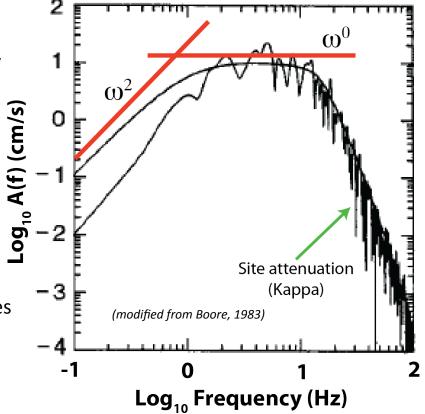
C_{ii} radiation scale factor

S_i(f) subfault source amplitude spectrum

G_{ii}(f) path term

P(f) site attenuation term

W_i*(f) complex spectrum of windowed time series



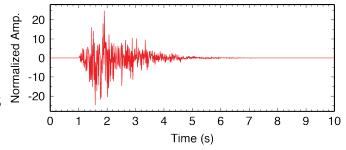




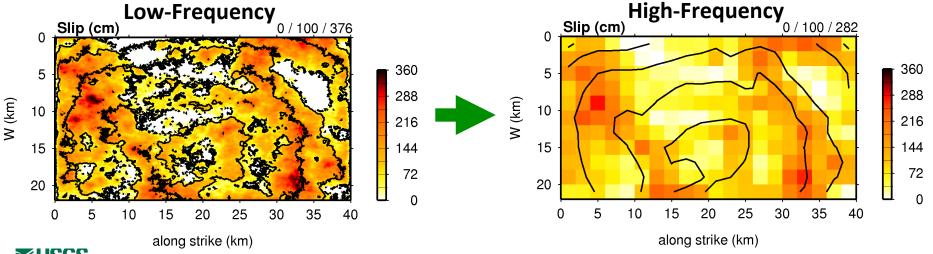
What does the Stochastic Phasing of W_i*(f) Represent?

The stochastic simulation approach, as originally developed by Hanks and McGuire (1981) and Boore (1983), was designed to match the statistical properties of observed high frequency ground motions using a simple, far-field model of the radiated amplitude spectra.

In this model, the stochastic phasing represents the **unmodeled details** of both the rupture process and scattering effects along the propagation path. Thus, there is no explicit slip-rate function in this approach and all of the detailed phasing effects from rupture across a subfault are incorporated stochastically within $W_i^*(f)$.



Since these features are represented stochastically, we do not require the fine spatial and temporal resolution of the kinematic rupture described earlier. In our HF implementation, we downsample the rupture to subfaults with dimensions of 1-2 km. We find 2 km X 2 km subfaults work well for WUS, and 1 km X 1 km work well for ENA (slight adjustments allowed to match overall fault dimensions).



Combining the Low- and High-Frequency Responses

Match Filtering

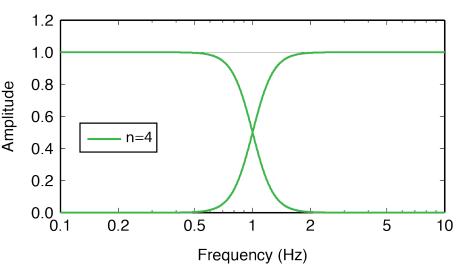
All hybrid simulation methods use some process of filtering and summation to combine the separate lowand high-frequency responses into a full broadband time series. There are a variety of approaches for this, each with its advantages and drawbacks. The basic assumption of all approaches is that the amplitude and phasing of the individual LF and HF responses are compatible across the cross-over frequency.

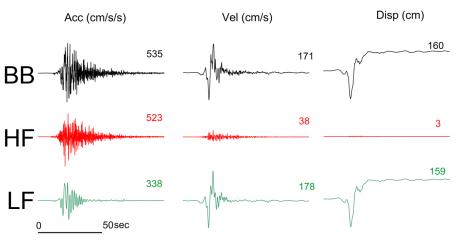
In our implementation, we use a set of "matched" 4^{th} order zero-phase Butterworth filters each with a corner frequency at f_m =1 Hz. A high-pass filter is applied to the HF response, and a low-pass filter is applied to the LF response. The filters sum to unity across all frequencies.

$$LP(f) = [1 + (f/f_m)^{2n}]^{-1} \qquad HP(f) = [1 + (f_m/f)^{2n}]^{-1}$$

The filtering is done in the time domain and the broadband response is obtained by summing the filtered results:

$$a_{BB}(t) = Ip(t)*a_{LF}(t) + hp(t)*a_{HF}(t)$$









Refinements to the Graves-Pitarka Broadband Simulation Method

Graves and Pitarka (2015)

- Addition of deep "weak-zone" to rupture characterization
- Perturbation of correlation structure for rise-time and rupture speed
- Extension of methodology to Eastern North America (ENA)



Hybrid approaches have traditionally been labeled using the following nomenclature:

- Low Frequency = Determinstic
- High frequency = Stochastic

Strictly speaking, the above classification is not correct because both approaches utilize deterministic and stochastic features. For example, the slip distribution used in the low-frequency simulation is generated using a stochastic model. Likewise, travel time and impedance effects in the high-frequency simulation are purely deterministic.

A more accurate representation is:

- Low Frequency = Comprehensive Theoretical Basis
- High frequency = Simplified Theoretical Basis



Complex Spectrum of Windowed Time Series W_i*(f)

$$A_{i}(f) = \sum_{j=1,M} C_{ij} S_{i}(f) G_{ij}(f) P(f) W_{i}^{*}(f)$$

This term incorporates all of the phasing information for the high frequency simulation.

Following Boore (1983) it is derived by first taking a windowed time sequence of band-limited random white Gaussian noise ("stochastic") and normalizing to have zero mean and scaling the variance to have unit spectral amplitude on average.

Our implementation uses the envelope function from Saragoni and Hart (1974) to shape the time sequence and set the peak of the envelope at the direct S-wave arrival time.

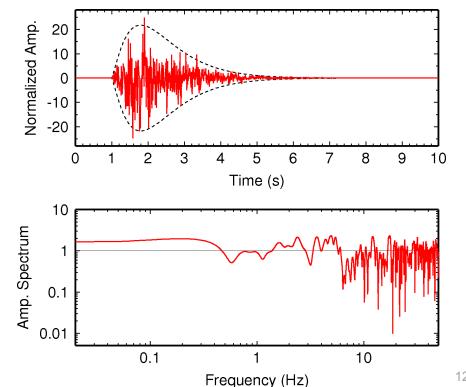
The duration of the windowed time sequence for the ith subfault is given by

$$T_{di} = f_{ci}^{-1} + c_T r_{ij}$$

where c_{τ} is set to 0.07 for WUS and 0.1 for ENA.

This term also incorporates the time delay for rupture propagation across the fault as well as the travel time for the particular ray being considered.

Once constructed in the time domain, this term is transformed into a complex valued frequency domain function that is then combined with the other amplitude shaping terms.





Kinematic Rupture Generator

Required inputs (SRC file)

- Magnitude
- Fault length (km)
- Subfault dimension along strike
- Fault width (km)
- Subfault dimension down-dip
- Latitude of top center point
- Longitude of top center point
- Depth to top of fault
- Hypocenter location along strike from top center (km)
- Hypocenter down-dip from top center (km)
- Strike
- Dip
- Rake (average)
- Seed for random number generator
- Time step for slip-rate function

Recommend using Leonard (2010) Magnitude-Area and Magnitude-Length scaling relations for GP method.



Kinematic Rupture Generator

Sample SRC file:

```
MAGNITUDE = 6.94
FAULT_LENGTH = 40.0
DLEN = 0.1
FAULT_WIDTH = 22.0
DWTD = 0.1
LAT_TOP_CENTER = 37.0789
LON_TOP_CENTER = -121.8410
DEPTH_TO_TOP = 0.0
HYPO\_ALONG\_STK = 0.0
HYPO DOWN DIP = 14.75
STRIKE = 128
DIP = 70
RAKE = 136
SEED = 1343642
DT = 0.1
```



1. Slip Distribution

- Begin with uniform slip having mild taper at edges.
- Transform to wavenumber domain and use Mai and Beroza (2002) spatial correlation functions with random phasing to filter wavenumber spectrum.

$$A(k_s, k_d) = [a_s a_d / (1 + K^2)^{H+1}]^{1/2} \qquad (H = 0.75)$$

$$K^2 = a_s^2 k_s^2 + a_d^2 k_d^2$$

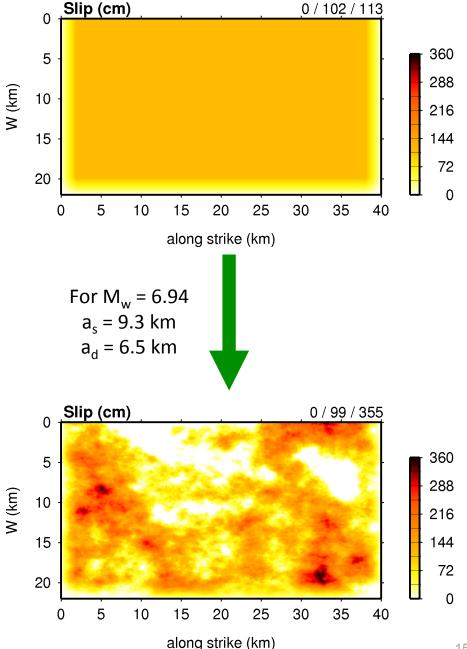
$$log_{10} a_s = M_w / 2.0 - 2.5$$

$$log_{10} a_d = M_w / 3.0 - 1.5$$

 Transform back to spatial domain and scale standard deviation of slip to be 85% of average:

$$\sigma_{s}$$
 = 0.85 \cdot D_{avg}

with D_{avg} scaled to give desired M_w





2. Rupture Initiation Time

First, compute background value T_B

$$T_B = r_{path} / V_r$$

WUS: $V_r = 80\%$ local V_s depth > 8 km = 56% local V_s depth < 5 km (linear transition between 5-8 km)

ENA: $V_r = 85\%$ local V_s all depths

Reduction of V_r above 5 km depth for WUS represents velocity strengthening behavior in weaker near-surface material

Next, add timing adjustment that correlates with local slip

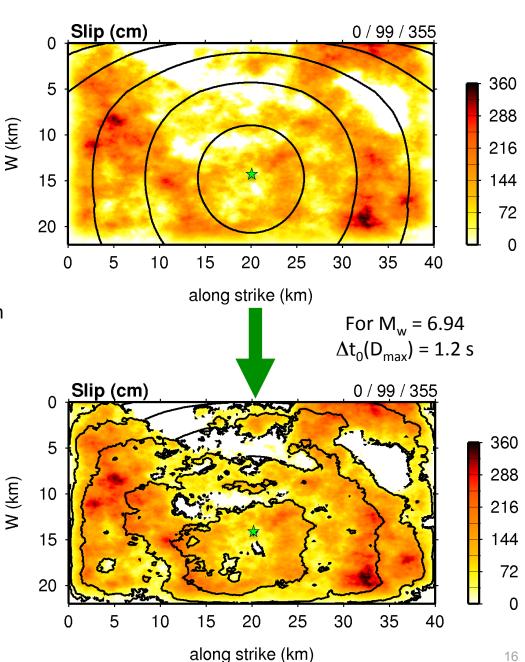
$$T_{i} = T_{B} - \Delta t_{0}(D_{i})$$

 Δt_0 scales with slip amount of ith subfault (D_i) to accelerate or decelerate rupture

$$\Delta t_0(D_{avg}) = 0$$

$$\Delta t_0(D_{max}) = 1.8 \times 10^{-9} \cdot M_0^{1/3}$$



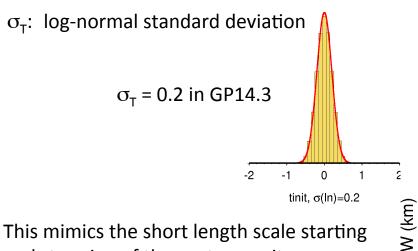


2. Rupture Initiation Time

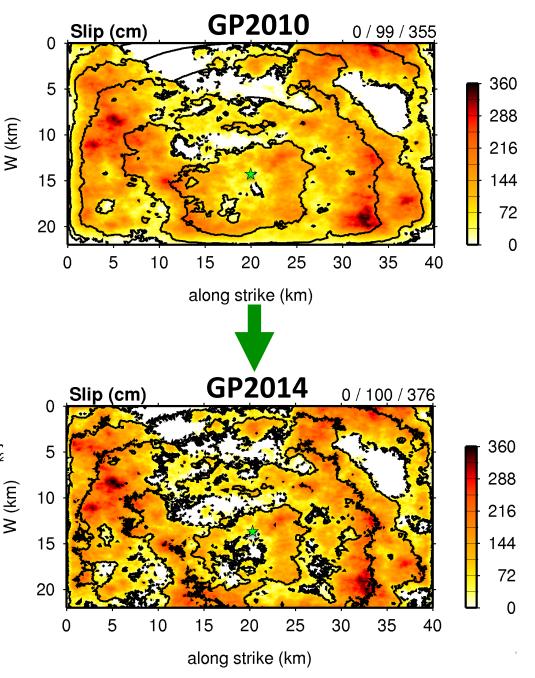
GP2014 adds random perturbations to timing adjustment so it is no longer correlated 1:1 with local slip

$$T_i = T_B - \Delta t_0(D_i) \exp(\epsilon \sigma_T)$$

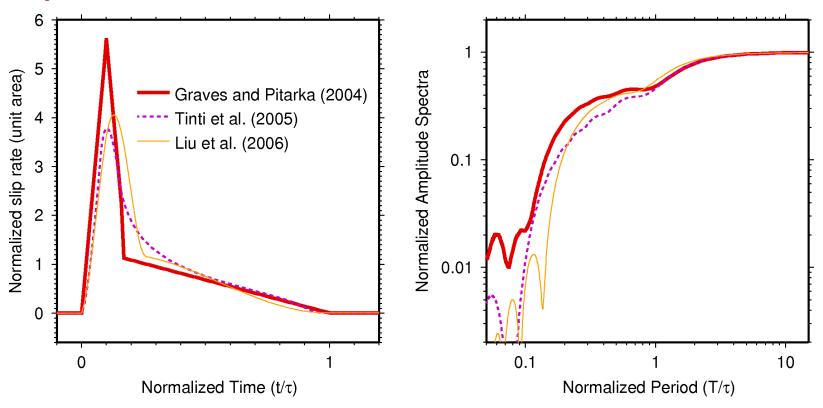
ε: random number selected from standard normal distribution (mean of zero, standard deviation of one)



This mimics the short length scale starting and stopping of the rupture as it propagates across the fault, possibly due to geometric complexities and/or stress heterogeneities







- We have considered several slip-rate functions in the GP methodology. All are "Kostrov-like" meaning they have a sharp rise to the maximum value followed by a lower amplitude, longer period tail.
- In our methodology, the rise time (τ) is the total length of the slip-rate function
- At periods around τ or longer there is little difference in these functions
- Currently, we use the slip-rate function based on Liu et al (2006)



Subfault rise time (τ_i) scales with square root of local slip (D_i)

WUS:
$$\tau_i = k \cdot D_i^{1/2}$$
 depth > 8 km
= $2 \cdot k \cdot D_i^{1/2}$ depth < 5 km
(linear transition between 5-8 km)

ENA:
$$\tau_i = k \cdot D_i^{1/2}$$
 all depths

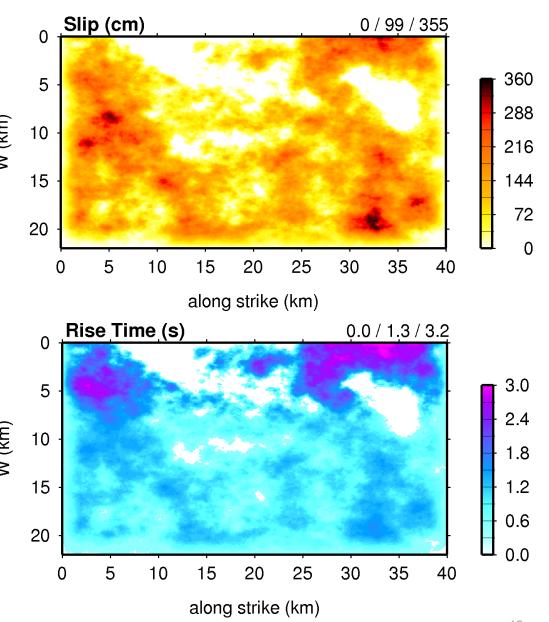
Lengthening of τ_i above 5 km depth in WUS represents velocity strengthening behavior in weaker near-surface material

The constant (k) set so average rise time across entire fault is given by the relation:

$$\tau_{A} = \alpha_{T} \cdot c_{1} \times 10^{-9} \cdot M_{o}^{1/3}$$

where c_1 =1.45 in WUS and 2.20 in ENA, and α_T is a mechanism dependent scaling factor (next slide).

The scaling with square root of slip represents a balance between constant rise time at one extreme and constant slip-rate at the other.





GP2014 modifications to rise time:

• The factor α_T scales the rise time as a function of fault dip and rake. The idea is that rupture in compressive regimes (e.g., blind thrust) produces shorter rise times.

$$\begin{split} \alpha_{\text{T}} &= [1 + \text{F}_{\text{D}} \, \text{F}_{\text{R}} \, c_{\alpha}]^{\text{-}1} \\ \text{F}_{\text{D}} &= 1 - (\delta - 45^{\circ})/45^{\circ}, \qquad 45^{\circ} < \delta \leq 90^{\circ} \\ &= 1, \qquad \delta \leq 45^{\circ} \\ \text{F}_{\text{R}} &= 1 - |\lambda - 90^{\circ}|/90^{\circ}, \qquad 0^{\circ} \leq \lambda \leq 180^{\circ} \\ &= 0, \qquad \text{otherwise} \\ c_{\alpha} &= 0.1 \end{split}$$

$$\min \alpha_{\text{T}} &= 0.91 \qquad \text{for } \delta \leq 45^{\circ} \text{ and } \lambda = 90^{\circ} \text{ (shallow thrust)} \\ \max \alpha_{\text{T}} &= 1.0 \qquad \text{for } \delta = 90^{\circ} \text{ (vertical fault)}, \lambda \leq 0^{\circ} \text{ (strike-slip/normal)} \end{split}$$

- For WUS, impose weak zone along the deeper portion of fault with factor of 2 increase in rise time from 15-18 km. Represents transition from unstable to stable sliding in midcrust (e.g., Scholz, 1998)
- Add perturbations to rise time so it is no longer correlated 1:1 with square root of local slip. Mimics short length scale variations possibly due to geometric complexities and/or stress heterogeneities

$$\tau_{Ri} = \tau_i \exp(\epsilon \sigma_{\tau})$$

 ϵ is a random number selected from standard normal distribution σ_{τ} is the log-normal standard deviation

distribution
$$\sigma_{\tau} = 0.5 \text{ in GP14.3}$$

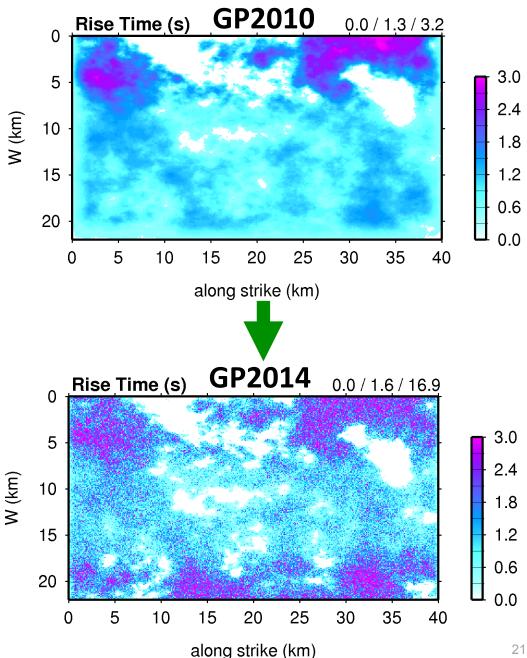
$$\sigma_{\tau} = 0.5 \text{ in GP14.3}$$

$$\sigma_{\tau} = 0.5 \text{ in GP14.3}$$



GP2014 modifications to rise time:

Goal is to provide a smoother transition between the low- and high-frequency simulation approaches through the addition of stochastic perturbations.





4. Slip Direction (Rake)

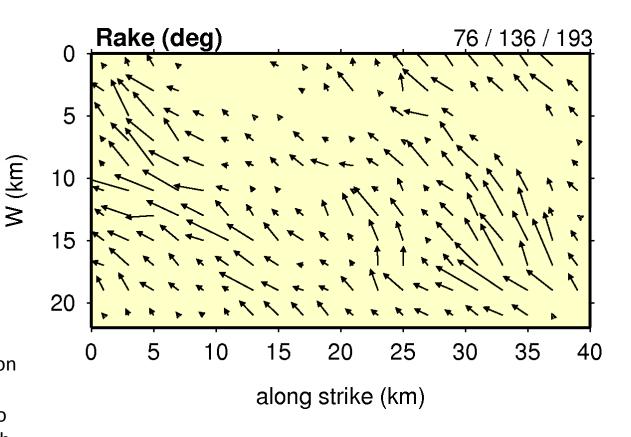
Subfault rake (λ_i) is given by average value (λ_0) plus random perturbations:

$$\lambda_{i} = \lambda_{o} + \varepsilon$$

range: $-60^{\circ} < \varepsilon < 60^{\circ}$

standard deviation: $\sigma_{\epsilon} = 15^{\circ}$

Random perturbations of rake follow spatial distribution given by Mai and Beroza (2002) wavenumber correlation structure (roughly K⁻² falloff). Uses a different seed than slip distribution so rake variations are not correlated with slip variations.





Green's Functions Table

The listing below shows the distances (1st block) and depths (2nd block) for a set of pre-computed WUS GFs.

Note that the sampling is not constant. It is finer at close distance and near the surface, and then increases with increasing distance and depth.

140						
0.1000	0.2000	0.4000	0.6000	0.8000	1.0000	
2.0000	3.0000	4.0000	5.0000	6.0000	7.0000	
8.0000	9.0000	10.0000	11.0000	12.0000	13.0000	
14.0000	15.0000	16.0000	17.0000	18.0000	19.0000	
20.0000	22.0000	24.0000	26.0000	28.0000	30.0000	
32.0000	34.0000	36.0000	38.0000	40.0000	42.0000	distance
44.0000	46.0000	48.0000	50.0000	52.0000	54.0000	block
56.0000	58.0000	60.0000	64.0000	68.0000	72.0000	
•						
•						
105 0000	470 0000	475 0000	100 000	405 0000	100 000	
465.0000	470.0000	475.0000	480.0000	485.0000	490.0000	
495.0000	500.0000					
50						
0.1000	0.2000	0.4000	0.6000	0.8000	1.0000	
1.2500	1.5000	1.7500	2.0000	2.2500	2.5000	
2.7500	3.0000	3.2500	3.5000	3.7500	4.0000	donth
4.5000	5.0000	5.5000	6.0000	6.5000	7.0000	depth
7.5000	8.0000	8.5000	9.0000	9.5000	10.0000	block
11.0000	12.0000	13.0000	14.0000	15.0000	16.0000	
17.0000	18.0000	19.0000	20.0000	21.0000	22.0000	
23.0000	24.0000	25.0000	26.0000	27.0000	28.0000	
29.0000	30.0000					
						┙



Green's Functions Interpolation

For each station in the LF simulation, the code loops over all the subfaults sequentially and sums the individual responses to obtain the total response. The GFs needed for each subfault are determined by the depth of the subfault and the distance from the subfault to the station, denoted here by d_E and r_E .

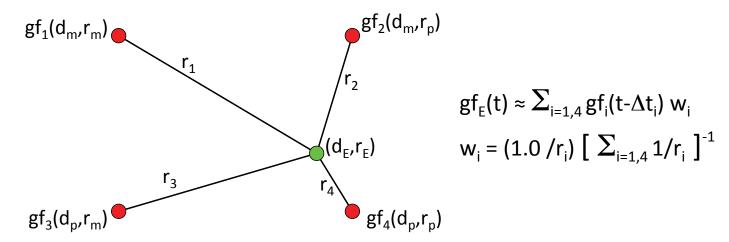
The code then looks up in the table to find the 2 closest depths and distances that bracket the desired location, i.e.,

$$d_{m} \le d_{E} \le d_{p}$$

$$r_{\rm m} \le r_{\rm E} \le r_{\rm p}$$

If there is an exact match to both the depth and distance, then that GF is selected and used in the simulation.

If an exact match is not found, then interpolation is done using weights given by the inverse distance to the exact point and by applying time shifts to align the S arrival.





Radiation Scale Factor C_{ij}

$$A_{i}(f) = \sum_{j=1,M} C_{ij} S_{i}(f) G_{ij}(f) P(f) W_{i}^{*}(f)$$

The radiation scale factor is given by

$$C_{ij} = F_s R_{Pij} / (4 \pi \rho_i \beta_i^3)$$

which represents the radiation for far-field S waves (Aki and Richards, 1980).

 $F_s = 2$ accounts for free surface amplification (assumes vertical incidence S-waves).

 R_{Pij} is a conically averaged radiation pattern term spanning a range of ±45° in slip mechanism and take-off angle for the jth ray.

 ρ_i and β_i are the density and shear-wave velocity at the center of the ith subfault.



Subfault Source Amplitude Spectrum S_i(f)

$$A_{i}(f) = \sum_{j=1,M} C_{ij} S_{i}(f) G_{ij}(f) P(f) W_{i}^{*}(f)$$

The subfault source amplitude spectrum is given by

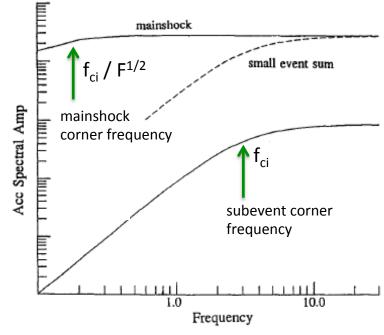
$$S_i(f) = m_i f^2 [1 + F \cdot f^2 / f_{ci}^2]^{-1}$$

which represents the ω^2 source radiation.

 $m_i = d_i \mu_i a_i$ is the seismic moment release of the ith subfault with d_i , μ_i , a_i being the slip, rigidity and area of the ith subfault.

 $f_{ci} = c_0 V_{Ri} / (\alpha_T \pi dl)$ is the subfault (subevent) corner frequency with V_{Ri} the local rupture speed (including shallow weak zone) and c_0 is set to 2.0 for WUS and 1.32 for ENA.

F = $\rm M_o$ / (N $\sigma_{\rm p}$ dl³) is a factor introduced by Frankel (1995), which scales the subfault corner frequency to that of the mainshock and ensures the total moment of the summed subfaults is the same as the mainshock moment $\rm M_o$. N is the total number of subfaults, $\sigma_{\rm p}$ is the Brune stress parameter (set 50 bars for WUS, 100 bars for ENA), and dl is the average subfault dimension.





Path Term G_{ij}(f)

$$A_{i}(f) = \sum_{j=1,M} C_{ij} S_{i}(f) G_{ij}(f) P(f) W_{i}^{*}(f)$$

The path term is given by

$$G_{ij}(f) = \{I_i(f) / r_{ij}\} \exp \{-\pi f^{1-x} \sum_{k=1,L} t_{ijk} / q_k \}$$

which represents impedance, geometric spreading and attenuation effects along the propagation path.

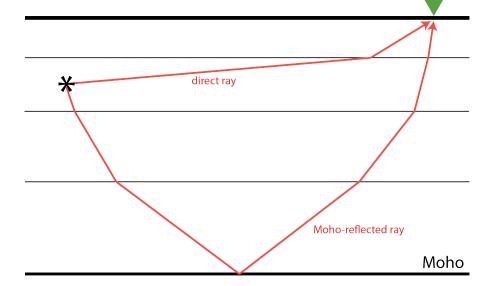
I_i(f) represents gross impedance effects from the ith subfault to the ground surface computed using quarter wavelength theory for the prescribed 1D velocity structure (Boore and Joyner, 1997).

 r_{ij} is the total path length of the j^{th} ray from the i^{th} subfault to the receiver.

exp{ \square } models anelasticity and scattering via a travel-time weighted average of the Q values for each of the velocity layers (Ou and Herrmann, 1990). The assumed frequency dependence is of the form Q(f)=Q₀f^x. The summation over k=1,L represents all of the ray path segments through the layers of the 1D velocity model, with t_{ijk} and q_k being the travel-time of the particular ray segment and Q value, respectively, within each velocity layer k.

WUS: x=0.6, $Q_0 \approx 120$

ENA: x=0.45, $Q_0 \approx 500$





Site Attenuation Term P(f)

$$A_{i}(f) = \sum_{j=1,M} C_{ij} S_{i}(f) G_{ij}(f) P(f) W_{i}^{*}(f)$$

The site attenuation term is given by

$$P(f) = \exp \left[-\pi \kappa_0 f\right]$$

which models near site high-frequency spectral decay based on the Kappa model introduced by Anderson and Hough (1984). We set κ_0 = 0.04 for WUS and κ_0 = 0.015 for ENA.



Combining the Low- and High-Frequency Responses

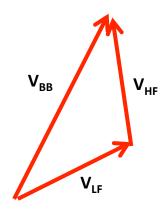
Some Caveats on Combining Hybrid Simulations

Generally, the separate low- and high-frequency simulations will not be exactly compatible around the cross-over frequency. In particular, any differences in phasing will cause a reduction of the summed amplitude response near the cross-over frequency.

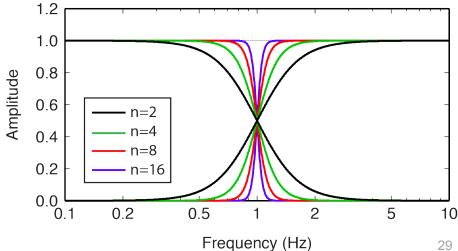
This occurs because the combination is a vector summation over both amplitude and phase. Any difference in phase between the LF and HF portions will act to reduce the combined BB amplitude.

$$\left| \left| V_{BB} \right| \le \left| \left| V_{LF} \right| + \left| V_{HF} \right|$$

The only case where the amplitude of the sum is not reduced is when the phase of the LF and HF portions are identical (which does not generally occur).

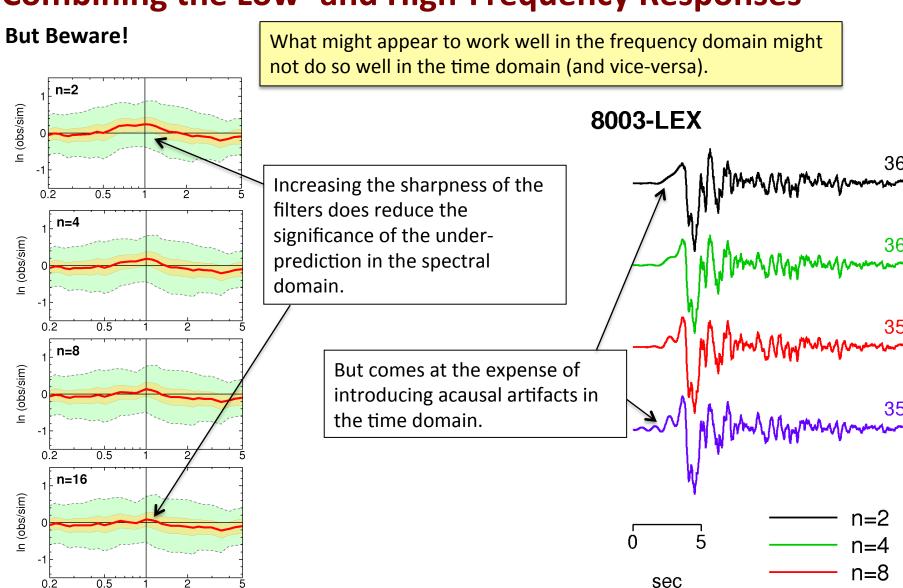


One way to minimize the impact of this effect is to make the filters as sharp as possible, for example, by increasing the order of the Butterworth operators. This doesn't eliminate the effect, but it concentrates it into a narrow frequency band.





Combining the Low- and High-Frequency Responses





Period (sec)

n=16

Research Needs

- Improve kinematic rupture characterization with guidance from rupture dynamics
- Improve models of non-linear response, both near fault and near surface
- Push "low-frequency" approach to higher frequencies
 - Ultimate goal is to eliminate need for "hybrid" approach; that is, develop a unified approach applicable to broad frequency range



Hidden Parameters

The current configuration of the BBP allows the user to specify some general parameters describing the source, station locations and Green's function. However, many of the underlying codes can accept a number of parameters that can be changed in the simulations. These parameters are "hidden" in the sense they cannot be changed using the standard BBP interface.

Modifying these parameters requires editing the python scripts that form the basis of the BBP workflow. Time permitting, this process will be discussed during the afternoon exercises.

For example, some of the relevant parameters in the GP method that a curious user might be interested in adjusting are:

Rupture Characterization:

rvfrac -specifies the background rupture speed as a fraction of local V_s (0.8 WUS, 0.85

ENA)

risetime_coef -gives the scaling of average risetime with seismic moment (1.45 WUS, 2.20 ENA)

High Frequency Simulation:

-specifies the background rupture speed as a fraction of local V_s (0.8 WUS, 0.85

ENA)

EXTRA FCFAC -gives the scaling of average corner frequency (0.0 WUS, -0.34 ENA)

 $(f_c = f_{c0}^*[1 + EXTRA_FCFAC])$



